

# Spectral Dimension of kappa space-time

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XXI DAE-BRNS HEP SYMPOSIUM - 2014, IIT Guwahati

09 December, 2014

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# Introduction

- Recent studies of space-time structure at high energies show the possibility of a dimensional flow.
- The effective dimension of a quantum space-time need not be same as the topological dimension.
- At high energy scale, space-time appears to be fuzzy due to quantum effects. The blurriness of the space-time screens some dimension, reducing the number of accessible dimensions.
- This gives us a hope for a renormalizable theory of gravity.

# What is spectral dimension?

- One can use diffusion process to measure the effective dimension, which is termed as spectral dimension.
- Intuitively : Corresponds to the dimension that a random walker 'experiences' in a diffusion process.
- Spectral dimension is known to change as one probes at high energies.

# Concepts of Spectral dimension

- **Diffusion equation**

$$\frac{\partial}{\partial \sigma} U(x, y; \sigma) = \mathcal{L}U(x, y; \sigma)$$

- $\sigma$  is a fictitious diffusion time.
- $\mathcal{L}$  is the Laplace operator.
- $U(x, y; \sigma)$  is the heat kernel representing the probability density of diffusion from  $x$  to  $y$  in diffusion time  $\sigma$ .

- **Return probability**

$$P_g(\sigma) = \frac{\int d^n x \sqrt{\det g_{\mu\nu}} U(x, x; \sigma)}{\int d^n x \sqrt{\det g_{\mu\nu}}}$$

- Logarithmic derivative of  $P_g(\sigma)$  will give the **spectral dimension**.

$$D_s = -2 \frac{\partial \ln P(\sigma)}{\partial \ln \sigma}$$

# Classical flat space-time

- Diffusion equation associated with a n-dimensional Euclidean geometry

$$\frac{\partial}{\partial \sigma} U(x, y; \sigma) = \nabla^2 U(x, y; \sigma)$$

Initial condition

$$U(x, y; 0) = \delta^n(x - y)$$

- Heat kernel

$$U(x, y; \sigma) = \frac{1}{(4\pi\sigma)^{\frac{n}{2}}} e^{-\frac{(x-y)^2}{4\sigma}}$$

- **Return probability**

$$P_g(\sigma) = \frac{\int d^n x \sqrt{\det g_{\mu\nu}} U(x, x; \sigma)}{\int d^n x \sqrt{\det g_{\mu\nu}}}$$

$$P_g(\sigma) = (4\pi\sigma)^{-\frac{n}{2}}$$

- **Spectral dimension**

$$D_s = n$$

- The spectral dimension is same as that of topological dimension  $n$ .
- It is independent of  $\sigma$ .

# Non-commutative space-time

- Concept of non-commutative (NC) space-time was introduced by H. Snyder to deal with the UV divergence in QFT.
- Different approaches to study the microscopic structure of space-time, introduced the notion of a minimum length scale (S. Doplicher, K. Fredenhagen, J. E. Roberts; Comm. math. phys. 172,1995).
- Space-time points will be replaced by cell and points become fuzzy.
- NC is a possible way to capture the structure of space-time at microscopic level.
- In recent times, NC space-time were brought back by developments in string theory and quantum gravity.



## Kappa deformed Euclidean space

- Coordinates obey Lie-Algebraic type of commutation relation.

$$[\hat{x}_n, \hat{x}_i] = ia\hat{x}_i \quad [\hat{x}_i, \hat{x}_j] = 0$$

where  $a$  is the deformation parameter.

- Two approaches:
  - Work with non-commutative functions defined on NC space-time.
  - Map function of NC coordinate to function on commutative space-time and use these non-local functions to setup the models (S. Meljanac and M. Stojic, Eur. Phys. J. C47 (2006) 531).

- Realization of the NC coordinates can be obtained in terms of commutative coordinates and its derivative by demanding linearity in  $x_\mu$  and  $\partial_\mu$ .

$$\begin{aligned}\hat{x}_i &= x_i \varphi(A) \\ \hat{x}_n &= x_n \psi(A) + i a x_k \partial_k \gamma(A)\end{aligned}$$

where  $A = ia\partial_n$

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$$[M_{\mu\nu}, M_{\lambda\rho}] = \delta_{\nu\lambda} M_{\mu\rho} - \delta_{\mu\lambda} M_{\nu\rho} - \delta_{\nu\rho} M_{\mu\lambda} + \delta_{\mu\rho} M_{\nu\lambda}$$

- Undeformed Poincare algebra with modified generators.

- Derivatives  $D_\mu$  which transform as a 4-vector under  $\kappa$ -undeformed Poincare algebra

$$D_i = \partial_i \frac{e^{-A}}{\varphi}, \quad D_n = \partial_n \frac{\sinh A}{A} + ia \nabla^2 \frac{e^{-A}}{2\varphi^2}$$

- Casimir of the undeformed  $\kappa$ -poincare algebra

$$D_\mu D_\mu = \square \left( 1 - \frac{a^2}{4} \square \right)$$

$$\square = \nabla^2 \frac{e^{-A}}{\varphi^2} - \partial_n^2 \frac{2(1 - \cosh A)}{A^2}$$

- Expand up to first non-vanishing terms in 'a'

$$D_\mu D_\mu = \nabla^2 + \partial_n^2 - \frac{a^2}{3} \partial_n^4 - \frac{a^2}{2} \nabla^2 \partial_n^2 - \frac{a^2}{4} \nabla^4$$

## Spectral dimension

- Modified diffusion equation

$$\frac{\partial U}{\partial \sigma} = \square \left( 1 - \frac{a^2}{4} \square \right) U$$

- First non vanishing terms in 'a'

$$\frac{\partial U}{\partial \sigma} = \nabla_n^2 U - \frac{a^2}{3} \partial_n^4 U - \frac{a^2}{2} \nabla^2 \partial_n^2 U - \frac{a^2}{4} \nabla^4 U$$

- Perturbative solution

$$U = U_0 + aU_1 + a^2U_2$$

- The return probability

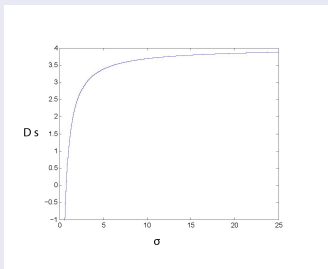
$$P_g(\sigma) = \frac{1}{(4\pi\sigma)^{\frac{n}{2}}} \left[ 1 + a\alpha + a^2\beta - a^2 \frac{(n+1)^2}{16\sigma} \right]$$

- Spectral dimension

$$D_s = n - (n+1)^2 \frac{a^2}{8\sigma}$$

$D_s$  vs  $\sigma$

$a = 1$     $n = 4$



## Conclusion

- For classical flat space-time the spectral dimension is same as the topological dimension.
- In case of NC space-time spectral dimension shows length scale dependence.

For a four dimensional space-time

- $\sigma \gg a^2$        $D_s \sim 4$
- $\sigma > a^2$        $D_s < 4$
- $\sigma \simeq a^2$        $D_s \simeq .875$
- $\sigma \ll a^2$        $D_s < 0$

## References

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**Thank You...**